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PIONEER 10/11 PROJECT

S-BAND OCCULTATION EXPERIMENT

REDUCED DATA

Explanatory Document

Submitted to the

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by

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Pioneer 10/11 Radio Occultation Experiment Data

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1. Introduction

The material submitted herewith consists of raw, intermediate, and reduced data from the Pioneer 10/11 S-Band Radio Occultation Experiment, covering data on Jupiter and Io. The following data are provided:

a) Digital raw data

Digital tapes only

b) Data Files

RPP, DIP1, DIP2, ATMOS (input)

Tapes, listings, and microfilm

ATMOS (output)

Listings and microfilm only

This material is described in detail in subsequent section.

2. Digital Data Tapes

Digital occultation data tapes are prepared by sampling analog spacecraft signal and time from analog recorded tapes. The sample rates are 60,000 samples/sec for Pioneer 10 and 10 and 40,000 samples/sec for Pioneer 11. These tapes contain a digital representation of recorded signals received from the spacecraft, the time of reception (UTC) and header information. They are 9-track, 800 BPI tapes with odd parity.

2.1 Format

Digital data tapes for Pioneer 10 and Pioneer 11 are in BCD mode. (odd-parity). Each record on the magnetic tape contains 4096 six-bit binary data words plus 9 six bit BCD words used to identify the time and one header word. The description of information in each record is shown in Table 1.

2.2 List of Digital Data Tapes

Promeer 10 and Pioneer 11 digital data tapes are listed in Table 2. All tapes are 9-track 800 BPI tapes. The tape number is in the second column and the start and stop times for the tape is given in the last column. Pioneer 10 and 11 and To tape, are listed separately.

LECEID					ar			*			
2005	FEATING.							.,	-,	.,	\$72.37 C2.07.02.02.
P 4	LONGITUDINAL EVEN	10	0	0	0	0	5	5	No.	85	NEADEA
旨	DATA SAUVLE BIT	P	0	0	0,	02	D ₃	Da	05	Do	Section of the sectio
5	STATION CODE	B	0	C	07	Da	Dg	010		11/2	Section Cont.
	10 - USS 14 01 - USS 4X	م	0	0	1/3	3	H ₅	1 .	A		energy appearance of the control of
	11 - DSS 5X	p	0	0	Ma	Me	ξ.	4	- 17	15	Service of the servic
[3	D - ANALOS	P	0	0		S3	Są		Sg		
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5 ₁ -8 ₇	(1) MELLISECCIO	San Manager Ma	0	0	US:		ra tulu ur	U ₂	G,	t! _a	197 C
ug-Mg2		2	Ū	>	-	4	1.3	W2			
Hy Fig	Com Mischesecond Burg	P	0	C	U.S	134	3	W2		0	•
P	CO PARITY BIT	p	0	0	W ₅	H ₂	ug	U ₂	S.C.	53	accom [*] cara wolo
		p	0	0	115	. 43		Į.		H _O	4005TH DATA 103.
		-	P	200	1	31	p.	g.	g:	0	
		- Advisoration to a partial prompting to appropriate company of the second seco								200 Dem (Ballandersche (Dunk) wiederstellen zur mit vorgen)	INTERDECOND 642
									MOT!		NOT USED IN CCA-1

Table 1: Data Record Format - Pioneer 10/11 Digital Occultation Data

Table: 2 Tapes (DSS 43)

PIONEER 10 EVENT	DIGITAL DATA TAPE NO.	TIME, UTC, Dec. 4, 1973
P10N (Entry)	x755	4:27:30 - 4:31:30 (Entry)
PlOX (Exit)	X681	5:28:00 - 5:32:00 (Exit)
PIONEER 11	DIGITAL DATA TAPE NO.	TIME, UTC, Dec. 3, 1974
Plinf (Entry FM)	X793	05:39:00 - 05:43:00 (Entry)
Plixf (Exit FM)	X658	06:22:30 - 06:26:30 (Exit)
PllXD (Exit Direct)	X641	06:22:30 - 06:26:30 (Exit)
PIONER 10 (Io)	DIGITAL DATA TAPE NO.	TIME, UTC, Dec. 4, 1973
IO1 (Entry)	X899	03:25:07 - 03:28:28
102 (Exit)	X944	03:28:30 - 03:31:29

3. Data Files

Pioneer 10 and Pioneer 11 data files are created by Processing Pioneer 10 and Pioneer 11 radio occultation data using the subset of occultation software.

A brief description of the software and the files is given below.

Occultation Software: consists of the Computer Programs RPP, DIP1, DIP2 and ATMOS. The functions of these programs are:

RPP removes drift and bias from frequency residuals supplied from open loop data

DIP1 computes refractive bending angle as a function of ray-asymptote distance

LTP2 computes refractivity as a function of radius from center of planet

a function of radius to center of planet

The Data Files described in Section 3.1 are Input/Output files obtained by executing the computer programs RPP, DIP1, DIP2, and ATMOS on the UNIVAC 1108.

The files are written using standard formatted FORTRAN I/O on the UNIVAC 1108 Exec-8 system. The format statements used are given in Section 3.3 according to the identifying number which is used as a label for each record in the detailed descriptions below.

Definitions and types of each variable are also given in Section 3.2. The type of a variable is either Integer, Real or Double Precision.

These files are written on tapes N936 and W058 (both 7-track, 800 BPI) using the tape processor of the 391*LIB. sub-routine library at JPL.

3.1 Occultation Data Tapes

Pioneer 10 and Pioneer 11 (Jupiter and Io) data files, obtained by executing the computer programs RPP, DIP1, DIP2, and ATMOS are discussed below.

3.1.1 Pioneer 10 and 11 (Jupiter) Data Files

The following data files are on tape N936 (7-track, 800 BPI) and are written using the TAPE processor. The sequence in which files are written is as follows:

File Sequence		
on N936	Descrip	tion
1.	P10 Entry OCEP	
2.	P10 Exit OCEP	
j.	P10 Exit OGEP	
4	Pl1 Entry OCEP	
5	Pll Exit OCEP	
6	P10 Entry TRAJ	
7	P10 Exit TRAJ	
8 .	Pll Entry TRAJ	
9	Pll Exit TRAJ	
10	P10 Entry RPP Output F	ile

11		P10 Entry DIP1 Output File
12	f =	P10 Entry DIP2 Output File
13	·	PlO Exit RPP Output File
14		PlO Exit DIP1 Output File
15		P10 Exit DIP2 Output File
16		P11 Exit RPP Output File -B1-
17		P11 Exit DIP1 Output File -B1-
18		Pll Exit DIP2 Output File -Bl-
19		Pll Exit RPP Output File -B2-
20		Pll Exit DIPl Output File -B2-
21		P11 Exit DIP2 Output File -B2-

3.1.2 Pionear 10 (Io) Data Files

The following are IO - Entry and Exit data files. These files are on tape W758 (7-trk, 800 BPI) and are written using TAPE processor. The first file on the tape is Trajectory file which is in a Type 66 format. All other files are in formatted FORTRAN I/O. The sequence in which files are written on tape is as follows:

File Sequence on W758 -	Description
1	TRAJ IO-entry and IO-exit T/66 Format
2	OCEP 10-entry
3	OCEP IO-exit
4	RPP output file IO-exit
5	DIP1 output file I0-exit
5	DIP2 output file IO-exit
7	File-13 RPP IO-entry
8	File-14 DIP1 IO-entry
9	File-15 DIP2 IO entry for ATMOS old F898 TRAJ
10	TRAJ IO-entry and IO-exit (FORTRAN Format)

3.2 Description of Data Files

The occultation software consists of the computer programs RPP, DIP1, DIP2, and ATMOS. These programs and the description of input/output files for them are given in the following sections.

3.2.1 Residual Processing Program (RPP)

The function of the Residual Processing Program (RPP) is to remove drift and bias from frequency residuals. The RPP program reads the RADIO DATA file produced by the OCEP (Occultation Editing Program) program which consists of time points and corresponding frequency residuals. RPP then computes a polynomial drift function for a user-specified time interval and subtracts this function from the data. RPP then computes the phase differences in cycles and write an output file consisting of doppler residuals (from which bias and drift has been removed) and amplitude to be used by the DIP1 program.

3.2.1.1 Input File

The input file to be used by RPP is the Radio Occultation data file produced by the OCEP Program. It consists of frequency residuals and power for each time points. This file is in standard formatted FORTRAN I/O and is described as follows. Each line corresponds to one record and the format statement for that record corresponds to the label (e.g. 5*)in section 3.3 definition and type of each variable in the file is also given.

FLTYPE	SCID	PASS	MODE		(Record 1)	*
XTX	DSS	IW	FBAND		(Record 2) 1	*
F1 .	F2				(Record 3) 2	×
SYNFRQ	SYNLO				(Record 4) 2	×
IIR	IDAY	TIME	SFREQ	SRES	(Record 5) 3	*
SPWR	XPWR	XFREQ	XRES		(Record 6) 4	×
•	•				2	
•		:	•			
IYR	IDAY	TIME	SFREQ	SRES	(Record n-1)] 3	÷
SPWR	XPWR	XFREQ	XRES		(Record n)	. ;
Whare,						
FLIYPE	-, -	f file. 1 and 3 for be		oop, 2 for o	pen (integs=)	
SCID	D spacecraft identification number					

PASS	orbit or revolution number	(integer)
MODE	1 for entry, 2 for exit	(integer)
XTR	transmitting station number	(integer)
DSS	receiving station number	(integer)
IW	1-way, 2-way or 3-way data	(integer)
FBAND	always 3	(integer)
F1	uplink frequency in hertz	(d. p.)
F2	downlink frequency in hertz	(d. p.)
SYNFRQ	closed loop synthesizer frequency	(d. p.)
SYNLO	open loop synthesizer frequency	(d. p.)
IYR	year of the sample	(integer)
IDAY	day of the sample	(integer)
TIME	time of sample in seconds past midnight	(real)
SFREQ	S-Band frequency in hertz	(d. p.)
SRES	S-Band frequency residuals in hertz	(d. p.)
SPWR	S-Band power in db.	(real)
XPWR	X-Band power in db.	(real)
XFREQ	X-Band frequency in hertz	(d. p.)
XRES	X-Band frequency residuals in hertz	(d. p.)

3.2.1.2 Output File:

RPP output file consists of doppler residuals (from which bias and drift has been removed) and phase for corresponding time points. The file is written with formatted FORTRAN I/O, and is described as follows:

FLTYPE	SCID	PASS	MODE		(Record 1)	1:
XTR	DSS	IW	FBAND	1	(Record 2)	1*
F1	F2				(Record 3)	2%
TIME	RESIDUAL	PHASE			(Record 4)	5®
•	•	•				
6	•	•			*	
TIME .	RESIDUAL	PHASE			(Record n)	5*
Where,						
FLTYPE			or closed lo	op, 2 for		
	open Loop	\cdot and 3 for	or both		(integer)	

SCID	spacecraft identification number	(integer)
PASS	orbit or revolution number	(integer)
MODE	1 for entry, 2 for exit	(integer)
XTR	transmitting station number	(integer)
DSS	receiving station number	(integer)
IW	1-way, 2-way, or 3-way data	(integer)
FBAND	always 3	(integer)
F1	uplink frequency in hertz	(d. p.)
F2	downlink frequency in hertz	(d. p.)
TIME	time of sample in seconds past midnight	(d. p.)
RESIDUAL	frequency sample in hertz	(d. p.)
PHASE	phase difference in cycles	(d. p.)

3.2.2 Data Inversion Program No. 1 (DIP1)

The function of DIP1 Program is to compute refractive bending angle, ray asymptote distance, and range from S/C to the center of the planet for each time point. Computation of the above variables are based on the assumptions of the oblateness in the Planet Jupiter. The exact shape of the planet Jupiter is computed by making use of a spherical harmonic representation of the gravity field of Jupiter and radius of curvature at point of ray tangency.

The input and output for DIP1 are described as follows:

3.2.2.1 Input files

Input files to DIP1 are the trajectory file and RPP output file. The RPP output file is described in Secting 3.2.1.2 on Page 7. The trajectory file is described as follows:

3.2.2.2 Trajectory file:

This file is a Trajectory file produced by the Trajectory link of the OCEP program and based on the DPTRAJ save and plot tape. The file is dynamically catalogued and assigned to the run. The file is written using formatted FORTRAN I/O, and the format is described as follows:

```
1 %
                                                               (Record 1)
                           PASS
                                       Ι
 FLTYPE
               SCID
                                                                             5*
                                                               (Record 2)
                           LITTM
 ETSCA
               ETMC
                                                               (Record 3)
                                                                             5*
                           S3
               S2
 SI
                                                      SEC(1)
                          DAY(1)
                                            MIN (1)
 TIME (1)
               YŘ(1)
                                    HR(1)
                                                                             6*
                                                               (Record 4)
 DPREC (1,1)
                          DPREC(1,3)
               DPREC(1,2)
                                                                             5*
                                                               (Record 5)
 DPREC (1,4)
                                                                             5×
                                                               (Record 6)
               DPREC (1,5) DPREC (1,6)
 DPREC(1,7)
                                                               (Record 7)
                                                                             5*
              DPREC (1.7) DPREC (1.8)
                                                               (Record 8)
                                                                            5*
              DPREC (1,11) DPREC (1,12)
TIME (N)
                          DAY (N)
                                    HR(N)
                                                               (Record M-4) 6*
                                              MIN(N) SEC(N)
DPREC(N,1)
                                                               (Record M-3) 5*
              DPREC(N,2) DPREC(N,3)
DPREC(N,4)
              DPREC(N,5) DPREC(N,6)
                                                               (Record M-2) 5*
DFMLC(N,7)
                                                              (Record M-1) 5*
              DPREC(N.8) DPREC(N.9)
                                                              (Record M)
                                                                            5%
              DPREC(N.11) DPREC(N.12)
Where,
               Type of file, 1 for closed loop, 2 for open loop
PLIYPE
               and 3 for both (INTEGER)
               Spacecraft Identification Number (INTEGER)
SCID
               Orbit or Revolution Number (INTEGER)
PASS
I
               1 for Entry, 2 for Exit (INTEGER)
               Time of closest approach - ET (sec/past Midnite 1950) (D.P.)
ETSCA
               Difference between Ephemeris time and Universal time/sec. (D.P.) .
ETHIC
               Time in Seconds (ET) from Jan. 1, 1950.0 for Nth Point (D.P.)
TIME (N)
               Year for the Nth Point (Real)
YR (N)
               Day number for Nth Point (Real)
DAY (N)
              Time at MIN minutes for Nth Point (Real)
MIN (N)
              Time at SEC - Seconds for Nth Point (Real)
SEC (N)
              1-way light time (D.P.)
LITTM
```

X, Y, Z Coordinates of Sun Position vector

relative to Jupiter (D.P.)

S1, S2, S3

DPREC (N,I)	I = 1, 2, 3
	X, Y, Z Coordinates of S/C Position vector relative
	to Jupiter for N th point (D.P.)
DPREC (N,I)	I = 4, 5, 6
	X, Y, Z Components of S/C Velocity vector realtive
	to Jupiter (D.P.) for N th Point
DPREC (N,I)	I = 7, 8, 9
	X, Y, Z Coordinates of Earth Position vector for
	N th Point (D.P.)
DPREC (N,I)	I = 10, 11, 12
	X, Y, Z Coordinates of Earth Velocity vector for
	N th Point (D.P.)
	N = 1, 2 - M Points

NOTE:

All above vectors are body centered and referenced to the Earth Equator of 1950.0 coordinate system

3.2.2.3 Cutput file:

DIP1 output file consists of ray-asymptote, bending angle, gravity and S/C range corresponding to each time point. The file is written with formatted FORTRAN I/O, and is described below.

FLTYPE	SCID PASS MODE	(Record 1) 1*
XTR	DSS IW FBAND	(Record 2) 1*
F1	F2	(Record 3) 2*
TIME	RD	(Record 4) 2*
BA	R	(Record 5) 2*
G	LAT	(Record 6) 2#
	•	
•.	•	•
•	•	
TIME	RD	(Record(n-2))2
BA	R	(Record (n-1))2
G	LAT	(Record n) 2*
Where,		
FLTYPE	Type of file. 1 for closed and 3 for both	loop, 2 for open loop (integer)
SCLD.	Spacecraft identification nu	umber (integer)

PASS	Orbit or revolution number	(integer)
MODE	1 for entry, 2 for exit	(integer)
XTR	Transmitting station number	(integer)
DSS	Receiving station number	(integer)
IW	1-way, 2-way or 3-way data	(integer)
FBAND	Always 3	(integer)
F1	Uplink frequency in hertz	(d. p.)
F2	Downlink frequency in hertz	(d. p.)
TIME	Time of sample in seconds past midnight	(d. p.)
G	Gravity Value	(d. p.)
LAT	Latitude of ray tangent in degrees	(d. p.)
RD	Asymptotic ray distance in Km	(d. p.)
BA	Refractive bending angle in radians	(d. p.)
R ,	Range from spacecraft to the center of the planet	(d. p.)

3.2.3 Data Inversion Program No. 2 (DIP2)

The function of the DIP2 program is to compute refractivity as a function of the radial distance to the center of the planet.

DIP2 reads a file of ray distance and corresponding bending angle

(among other parameters) produced by the DIP1 program and inverts this data
using the Abel Integral Transform to obtain refractivity versus radius.

3.2.3.1 Input file:

DIP2 input file is same as DIP1 output file. This file is in formatted FORTRAN I/O, and is described on Page 10, Section 3.2.2.3.

3.2.3.2 Output file:

The DIP2 output file is written with formatted FORTRAN I/O. It consists of radius, refractivity, gravity and latitude and is described as follows:

FLTYPE	SCID	PASS	MODE	(Record 1) 1*
XTR	DSS	IW	FBAND	(Record 2) 1*
F1	F2			(Record 3) 28
RADIUS	REF			(Record 4) 2*

GR	LATT	(Record	5) 2*	
•	•	•		
•	•	•		
•	•			
RADIUS	REF	(Record	(n-1))2	×
GR	LATT	(Record	n) 2	*
Where,				
FLTYPE	type of file. 1 for closed loop, 2 for open and 3 for both	100p	(intege	er)
SCID	spacecraft identification number		(intege	er)
PASS	orbit or revolution number		(intege	er)
MODE	1 for entry, 2 for exit		(intege	er)
XTR	transmitting station number		(intege	er)
DSS	receiving station number		(intege	er)
L'.J	1-way, 2-way or 3-way data		(intege	er)
FBAND	always 3		(intege	er)
21	uplink frequency in hertz		(d. p.)	N
. 1943	donwlink frequency in hertz		(d. p.)	
RADIUS	radial distance to the center of the planet :	in Km	(d. p.)	
REF	refractivity in N-units	(*)	(L. p.)	
GR	gravity value obtained from NALPHA		(d. p.)	
LIT	latitude of ray tangent in degrees		(c. p.)	

3.2.4 Atmospheric Parameter Program (ATMOS)

The function of the AMTOS program is to compute atmospheric information from the refractivity data produced by the DIP2 Program.

A detailed description of ATMOS can be found in the Mariner Mars 1971 occultation Atmospheric Parameter Program - Program document.

3.2.4.1 Input File

ATMOS input file is the same as DIP2 output file. This file is in formatted FORTRAN I/O and is described on Page 11, Section 3.2.3.2.

3.2.4.2 Output

The output of the ATMOS Program consists of hard-copy computer listings and plots containing the following information:

For all points:

RADIUS (km) REFRACTIVITY (n-units) with bias removed

For negative refractivity points:

RADIUS (km) REFRACTIVITY (n-units) ELECTRON DENSITY (cm⁻³) (these data are not reliable below main peak)

For all positive refractivity points:

RADIUS (km) REFRACTIVITY (n-units) TEMPERATURE (°K) PRESSURE (mb)

NUMBER DENSITY (cm⁻³) MASS DENSITY (gcm⁻³)

3.3 Formats

The following formats are referred to in the preceding:

- 1* FORMAT (4110)
- 2* FORMAT (2D26.18)
- 3* FORMAT (12,15,F12.4,2D26.18)
- 4* FORMAT (F9.4,F10.4,2D26.18)
- 5* FORMAT (3D26.18)
- 6* FORMAT (3D26.14,4X,4F7.1,F14.8)
- 7* FORMAT (4D26.18)

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