

#556

ISEE 1

PLASMASPHERIC ELEC. FLD. 3 SEC. AVG

77-102A-11D

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1. INTRODUCTION:

The documentation for this data set was originally on paper, kept in NSSDC's Data Set Catalogs (DSCs). The paper documentation in the Data Set Catalogs have been made into digital images, and then collected into a single PDF file for each Data Set Catalog. The inventory information in these DSCs is current as of July 1, 2004. This inventory information is now no longer maintained in the DSCs, but is now managed in the inventory part of the NSSDC information system. The information existing in the DSCs is now not needed for locating the data files, but we did not remove that inventory information.

The offline tape datasets have now been migrated from the original magnetic tape to Archival Information Packages (AIP's).

A prior restoration may have been done on data sets, if a requestor of this data set has questions; they should send an inquiry to the request office to see if additional information exists.

2. ERRATA/CHANGE LOG:

NOTE: Changes are made in a text box, and will show up that way when displayed on screen with a PDF reader.

When printing, special settings may be required to make the text box appear on the printed output.

Version	Date	Person	Page	Description of Change
01				
02				

3 LINKS TO RELEVANT INFORMATION IN THE ONLINE NSSDC
INFORMATION SYSTEM:

<http://nssdc.gsfc.nasa.gov/nmc/>

[NOTE: This link will take you to the main page of the NSSDC Master Catalog. There you will be able to perform searches to find additional information]

4. CATALOG MATERIALS:

- a. Associated Documents To find associated documents you will need to know the document ID number and then click here.
<http://nssdcftp.gsfc.nasa.gov/miscellaneous/documents/>

- b. Core Catalog Materials

ISEE 1

3-S AVGD PLASMASPHERIC ELEC. FIELD

77-102A-11D SPMS-00255

THIS DATA SET HAS BEEN RESTORED. THERE WAS ORIGINALLY ONE 9-TRACK, 1600 BPI TAPE WRITTEN IN EBCDIC. THERE IS ONE RESTORED TAPE WRITTEN IN ASCII. THE DR TAPE IS A 3480 CARTRIDGE AND THE DS TAPE IS 9-TRACK, 6250 BPI. THE TIME SPAN IS NOT CONTAINED ON THE TAPE AND CAN NOT BE VERIFIED. THE ORIGINAL TAPE WAS CREATED ON AN IBM 360 COMPUTER AND WAS RESTORED ON AN IBM 9021 COMPUTER. THE DR AND DS NUMBER ALONG WITH THE CORRESPONDING D NUMBER AND TIME SPAN IS AS FOLLOWS:

DR#	DS#	D#	FILES	TIME SPAN
DR005346	DS005346	D053321	160	12/02/77 - 11/30/78

REQ. AGENT

REQ. NO.

ACQ. AGENT

DEW

V0160

HKH

ISEE 1
PLASMASPHERIC ELEC. FLD. 3 SEC. AVG
77-102A-11D

This data set catalog consists of 1 data tape. The tape is 1600, bpi, 9 track, EBCDIC, with 160 files of data. The time span is not on the tape. The D and C numbers are as follows:

<u>D#</u>	<u>C#</u>	<u>TIME SPAN</u>	<u>FILES</u>
D-53321	C-22722	12/2/77 - 11/30/78	160

September 15, 1982

TO: 601/NSSDC/Kent Hills
FROM: 696/Computer Sciences Corporation
SUBJECT: 9 Track 1600 bpi Tape of the ISEE-1

Attached is a 9 track 1600 bpi tape of the ISEE-1 culled data sets for a 12 month period beginning in December 1977. In addition, there is a chronological listing of the files noting the start and finish times and whether the satellite was in-bound or out-bound. There are 166 files on the tape and the following is a list of their characteristics:

The recordlength is 55 bytes (the 55th byte is a control character) and the block size is 2750 bytes.

Format (3I2, 6F8.3, 1x)

	TIME								
	HH	MM	SS	L	MLT	E_n	E_E	E_R	E_A
Byte#	12, 34, 56,	7..14,	15..22,	23..30,	31..38,	39..46,	47..54		

where E_n and E_a are the electric field components in the equatorial plane and E_n and E_E are their ionospheric projections. A data description is attached.

Please contact me if you have any questions.

Cathie Meetre

Cathie Meetre (344-5403)

Attachments: 2

September 14, 1982

TO: 601/NSSDC/Kent Hills
FROM: 696/Electrodynamics Branch
SUBJECT: ISEE-1 Electric Field Data

Attached is (1) a data tape, (2) tape format, (3) listing of files and (4) a brief description of how the data was handled for ISEE-1 Electric Field data from the HPM instrument. The tape covers from December 1, 1977 to December 1, 1978 and includes all usable data in the plasmasphere (mostly inside L of 6). The listing of files gives each UT time span. We would like to submit this data to the NSSDC ISEE files. For more information contact N. C. Maynard.



James P. Heppner

Attachments: 4

HPM PLASMAPHERIC ELECTRIC FIELDS

The Electric Field Data

Electric field measurements were made from the ISEE-1 spacecraft using the double probe technique with long cylindrical sensors (Heppner et al., 1978). The baseline (separation between the midpoints of the sensors) of the double probe system of long wires was 179 m. From this single axis, measurements of both components of the electric field in the spin plane were possible every spin period (3.05s). Data along the instantaneous direction of the axis were taken 16 or 32 times per second depending on the data rate. The spin plane was approximately coincident with the solar ecliptic XY plane in this 23.2 Re apogee orbit.

In all data presented here the $\underline{v}_s \times \underline{B}$ electric field from the satellite velocity \underline{v}_s and the corotation electric field have been vectorially subtracted. Thus the measurements are transferred to a coordinate system corotating with an observer on the Earth. The magnetic field used in the $\underline{v}_s \times \underline{B}$ calculations was that measured on board the spacecraft (Russell, 1978).

To determine the electric field in the spin plane, data taken when either antenna was within 20° of the sun was culled in order to remove any errors caused by shadowing of a sensor or by reduced photoemission from a small sun angle, and the remaining data was fitted to a sine wave using a least squares technique. The solar ecliptic X and Y electric field components in inertial space were then determined from the sine wave fit, and the $\underline{v}_s \times \underline{B}$ and corotation subtractions were done to yield the electric field in the earth fixed coordinate system.

The third component of the electric field (solar ecliptic E_z) was calculated assuming that $\underline{E} \cdot \underline{B} = 0$. This extrapolation is good providing that the ratio of \underline{B}_{xy} to \underline{B}_z is low. The resulting vector electric field was then transformed into the radially outward and azimuthally eastward

components perpendicular to \underline{B} . For the plasmasphere studies these electric fields were then projected to the magnetic equatorial plane and to the ionosphere assuming that the magnetic field lines are equipotentials and that the magnetic field was a dipole. The projections were done on each data set to provide a common reference for comparison of data taken at different magnetic latitudes. When projected to the northern hemisphere ionosphere these components become the northward and eastward electric fields respectively. Note that the projection introduces errors from our imprecise knowledge of the magnetic field configuration. The original L values were calculated using the GSFC harmonic analyses of B with the Mead-Fairfield tail extension. We in turn have projected the values to the ionosphere assuming a symmetric dipole. Thus especially at the higher L values, both the projection factor and location will have some error. This was not considered to be critical to the results presented here.

The purpose of this study was to find the average electric field seen in the plasmasphere. Thus it was mandatory to remove any bad data that could affect the results. All data were scanned, and regions where the instrument was in its internal calibration cycle as well as regions where the Harvey instrument (Harvey et al., 1978) was transmitting (causing the electric field sensors to be shorted) were culled from the data set.

In regions of low density plasma ($< 1/\text{cm}^3$) the electric field instrument sees an apparent sunward directed electric field that is caused by the asymmetries of the sheath of photoelectrons around the spacecraft. The effect results from two causes. The dipole of the negatively charged sheath on the sunward side of the spacecraft and the positively charged spacecraft is one cause, and the other is the asymmetric collection by the sunward sensor of photoelectrons from the spacecraft, (see Cauffman and Maynard, 1975). It is hard to pinpoint exactly at what density this effect becomes significant (it is also a function of the temperature of the medium as well). As the plasma becomes more rarified, the first sign of asymmetrical photoelectron emission errors appears in data taken with the sensor axis within $\pm 20^\circ$ from the sun or at a near grazing angle (data which is not used in the least squares calculation to determine the electric field). In order to scrupulously avoid any data that might be the least

bit contaminated by this effect, all data were culled that had large bite outs where the antenna axis was directed toward the sun or that had a significant solar ecliptic E_x . This last criteria eliminated some good data where the convective electric field would be expected to be in that direction. A comparison of the point where culling started and plots of electron density from the Harvey instrument for several orbits (M. Echeto, private communication, 1980) showed that the data that were kept were in general from regions of electron densities greater than 30 to 50/cm³.

The ISEE-1 electric field data from 1 December 1977 to 1 December 1978 were scanned and culled with the above criteria to form the data base for this study. A second culling of the data base was done to remove data when the ratio of B_z to $|B|$ was 0.2 or less in order to prevent errors from the $\underline{E} \cdot \underline{B}$ extrapolation from seriously effecting the results.

ISEE-1 Culled data sets
in chronological order

December 1977 - November 1978

Prepared by C. Heette, C.S.C.
July 1982

December 77

Date	Start time			Stop time			Tape File #	I	O	N	Disc #
DEC 2, 77	4	54	50	5	30	0	1	✓			✓ 14
DEC 4, 77	13	45	02	14	56	26	2	✓			✓ 14
DEC 4, 77	15	58	25	16	29	55	3		✓		✓ 14
DEC 6, 77	23	0	1	23	59	55	4	✓			✓ 14
DEC 7, 77	0	0	7	0	19	44	5	✓			✓ 14
DEC 7, 77	1	24	26	2	04	55	6		✓		✓ 14
DEC 9, 77	10	48	35	11	08	38	7		✓		✓ 14
DEC 9, 77	7	40	1	9	9	16	8	✓			✓ 14
DEC 11, 77	18	12	06	19	09	59	9	✓			✓ 14
DEC 14, 77	3	42	40	4	24	59	10	✓			✓ 14
DEC 14, 77	5	32	12	5	37	49	11		✓		✓ 14
DEC 16, 77	14	55	28	15	30	59	12		✓		✓ 1
DEC 16, 77	11	38	36	13	34	57	13	✓			✓ 1
DEC 18, 77	21	48	06	23	09	10	14	✓			✓ 1
Dec 19, 77	0	24	06	0	54	58	15		✓		✓ 1
DEC 21, 77	9	44	44	9	57	53	16		✓		✓ 1
Dec 21, 77	7	40	06	8	14	59	17	✓			✓ 1
Dec 23, 77	19	13	07	19	33	06	18		✓		✓ 1
Dec 23, 77	16	04	07	17	59	03	19	✓			✓ 1
Dec 26, 77	2	13	7	3	13	9	20	✓			✓ 2
Dec 28, 77	13	52	03	14	19	55	21		✓		✓ 1
Dec 28, 77	10	41	7	12	40	55	22	✓			✓ 1
Dec 30, 77	20	22	06	21	51	50	23	✓			2
Dec 30, 77	23	14	38	23	29	58	24		✓		2

(23)

mislabelled on disc as DEC1278IC

Date	Start time			Stop time			Tape file #	I	O	N	Disc #	
	Hr	Min	Sec	Hr	Min	Sec						
Jan 2, 1978	6	32	00	6	59	59	25	✓			✓	2
Jan 2	8	40	43	8	46	35	26		✓		✓	2
Jan 4	18	8	16	18	18	8	27		✓		✓	2
Jan 4	16	19	02	16	43	55	28	✓			✓	2
Jan 9	12	52	11	12	59	58	29		✓		✓	2
Jan 9	10	22	26	11	28	42	30	✓			✓	2
Jan 11	19	38	01	20	49	58	31	✓			✓	2
Jan 11	22	12	14	22	18	58	32		✓		✓	2
Jan 14	4	57	41	5	15	53	33	✓			✓	2
Jan 16	13	20	03	15	24	59	34	✓			✓	2
Jan 18	23	49	0	23	59	51	35	✓			✓	2
Jan 19	0	0	9	0	46	39	36	✓			✓	2
Jan 21	10	59	18	12	1	52	38	✓			✓	3
Jan 21	8	15	1	10	7	59	37		✓		✓	3
Jan 23	21	29	50	21	37	26	39		✓		✓	3
Jan 23	17	15	02	19	19	09	40	✓			✓	3
Jan 26	5	39	08	5	52	07	41		✓		✓	3
Jan 28	15	4	2	15	59	58	42		✓		✓	3
Jan 28	12	38	02	14	20	0	43	✓			✓	3
Jan 30	22	55	1	23	13	14	44	✓			✓	3
Jan 31	20	27	1	0	53	54	45		✓		✓	3

Date archi '78	Start time						Tape file #	In	Out	N	Disc #
	HR	MIN	SEC	HR	MIN	SEC					
3	2	17	7	2	37	56	62		✓		10
5	9	39	6	10	59	59	63	✓			10
5	11	39	48	12	19	0	64		✓		10
7	19	40	7	20	14	58	65	✓			10
10	4	30	8	5	38	18	66	✓			10
10	6	26	8	6	59	58	67		✓		10
12	13	53	7	15	17	58	68	✓			17
12	15	49	7	16	39	58	69		✓		17
15	0	10	8	2	45	9	70	✓			17
17	10	34	7	11	18	26	71		✓		17
17	9	8	8	9	54	59	72	✓			17
29	7	42	7	8	46	57		✓			5
29	9	48	7	10	6	58	73		✓		14
31	18	46	6	19	9	58	74 ?		✓		14

May '78

Date	Start			Stop			Tape file #	In	Out	N
May 78	23	10:00	46	23	29	37	0 N914	✓		
8	8	22	8	9	12	20	92	✓		
11	TRY THIS AGAIN								✓	
13	3	24	14	3	53	52	93			✓
16	4	35	7	5	32	59	94		✓	
16	13	3	7	13	10	0	95			✓
18	8	0	25	8	8	1	96	✓		
23	8	56	9	9	19	43	97		✓	
23	2	30	7	2	42	50	98			✓
28	1	5	7	1	22	32	99	✓		
28	3	22	7	4	16	36	100		✓	

June '78

Date	Start			Stop			Tape file #	In	Out	N	
June '78											
1	22	35	25	22	34	58	101		✓		12
1	20	0	6	20	18	46	102	✓			12
4	5	30	8	25	49	58	103	✓			12
4	7	36	9	8	18	6	104		✓		12
9	2	6	9	2	39	57	105		✓		12
11	11	25	8	12	24	59	106		✓		10
11	9	43	22	9	53	8	107	✓			10
13	20	52	6	20	57	29	108		✓		10
16	5	28	39	5	34	59	109			✓	10
20	22	41	4	23	25	0	110	✓			10
21	0	57	9	1	42	43	111		✓		10
23	8	31	8	8	51	58	112	✓			10
23	9	30	40	9	36	48	113			✓	10
23	10	9	8	10	39	45	114		✓		10
28	3	4	9	3	37	18	115	✓			10

July 78

Date	Start			Stop			Tape file #	I	O	N	Disc #
July 78											
2	22	4	36	22	23	55	116	✓			16
2	23	34	7	23	59	58	117		✓		16
3	20	0	19	0	10	0	118		✓		16
5	9	2	15	9	29	58	119		✓		16
10	1	36	17	2	28	32	120	✓			16
12	10	56	18	11	53	24	121	✓			16
14	21	0	6	21	19	59	122	✓			11
17	5	0	7	5	10	28	123	✓			15
17	8	22	16	8	53	52	124		✓		11
19	15	20	9	15	53	52	125	✓			11
22	0	40	9	1	21	26	126	✓			11
22	2	16	6	3	29	27	127		✓		11
24	9	54	7	10	43	56	128	✓			11
26	19	0	6	19	14	58	129	✓			11
26	20	58	0	21	45	0	130		✓		11
29	7	8	9	7	39	57	131		✓		11
29	4	25	2	5	29	57	132	✓			11

Date	Start		Stop		Tape file #	I	O	N	Disc#	
Aug '78										
3	1	21	6	2	8	19		✓		15
X										
5								✓	←	15
10	5	6	7	6	4	58		✓		15
10	3	15	1	4	20	0	✓			15
14	23	47	9	23	54	58		✓		15
15	0	11	18	0	24	59		✓		15
17	7	8	8	7	34	58	✓			15
X							✓			15
19	18	27	08	18	36	37		✓		15
22	3	51	31	4	9	59		✓		15
22	1	4	8	3	9	59	✓			15

\$NOP
\$NOP ***** DO-1 *****
\$EXEC TPLIST BS

053321

12/2/77 - 11/30/78

INPUT PARAMETERS ARE: ED SR=1=1 1

TAPE NO.	1	FILE NO.	1										
RECORD	1	LENGTH	2750										
45450	4.211	11.998	-1.615	-6.426	-0.103	-0.744	45453	4.209	11.998	-1.175	-8.526	-0.0	
75	-0.987	45456	4.208	11.999	-0.357	-8.769	-0.023	-1.016	45459	4.206	12.001	-0.929	-
7.372	-0.059	-0.855	455 2	4.204	12.002	-1.516	-7.053	-0.097	-0.818	455 5	4.202	12.003	-
-2.365	-8.592	-0.151	-0.997	455 8	4.201	12.005	-4.494	-9.362	-0.268	-1.087	45511	4.	
199	12.006	-5.633	-7.372	-0.361	-0.857	45514	4.197	12.008	-5.523	-7.797	-0.354	-0.907	
45517	4.196	12.009	-5.427	-8.940	-0.346	-1.040	45520	4.194	12.010	-6.145	-9.902	-0.3	
95	-1.153	45523	4.192	12.012	-6.272	-10.529	-0.403	-1.227	45526	4.190	12.013	-3.707	-1
0.622	-1.239	-1.233	45529	4.189	12.015	-1.719	-9.983	-0.111	-1.164	45532	4.187	12.016	-
-2.693	-6.390	-0.173	-0.735	45535	4.185	12.017	-3.280	-6.554	-0.211	-0.766	45538	4.	
183	12.019	-5.230	-5.328	-0.337	-0.623	45541	4.182	12.020	-6.425	-1.579	-0.415	-0.185	
45544	4.180	12.022	-5.792	0.457	-0.374	0.053	45547	4.178	12.023	-6.723	0.142	-0.4	
35	0.017	45550	4.176	12.024	-8.317	0.919	-0.538	0.106	45554	4.175	12.026	-5.659	
0.524	-0.366	0.061	45557	4.173	12.027	-2.716	-2.289	-0.176	-0.269	456 0	4.171	12.029	
-3.617	-1.328	-0.234	-0.156	456 3	4.170	12.030	-1.448	2.112	-0.094	0.248	456 6	4.	
168	12.031	-0.537	5.952	-0.035	0.699	456 9	4.166	12.033	1.782	6.197	0.116	0.729	
45612	4.164	12.034	2.482	6.370	0.161	0.750	45615	4.163	12.036	2.340	9.257	0.1	
52	1.090	45618	4.161	12.037	1.096	9.866	0.071	1.162	45621	4.159	12.039	3.199	1
0.745	0.298	1.267	45624	4.157	12.040	-4.262	7.210	-0.278	0.851	45627	4.156	12.041	
-8.157	6.888	-0.525	0.613	45630	4.154	12.043	-8.007	8.214	-0.522	0.970	45633	4.	
152	12.044	-12.121	8.538	-0.791	1.009	45636	4.151	12.046	-12.512	10.305	-0.617	1.218	
45639	4.149	12.047	-12.581	11.351	-0.822	1.343	45642	4.147	12.049	-14.734	11.445	-0.9	
64	1.355	45645	4.145	12.050	-15.083	10.826	-0.987	1.283	45648	4.144	12.051	-14.506	1
0.656	-0.950	1.263	45651	4.142	12.053	-12.016	9.487	-0.788	1.126	45654	4.140	12.054	
-6.708	7.280	-0.440	0.864	45657	4.138	12.056	-4.883	6.909	-0.321	0.821	457 0	4.	
137	12.057	-5.911	6.747	-0.388	0.802	457 3	4.135	12.059	-6.318	7.734	-0.415	0.920	
457 6	4.133	12.060	-7.904	7.736	-0.520	0.921	45710	4.132	12.062	-9.480	5.179	-0.6	
24	0.617	45713	4.130	12.063	-9.727	4.427	-0.641	0.527	45716	4.128	12.065	-10.653	
3.731	-0.702	0.451	45719	4.126	12.066	-10.935	2.274	-0.721	0.271				

***** JOB DONE.
\$WEO LPS

\$s
\$ASS IN TD2 OUT MT4
\$EXE TPDUPC BS